

## Telescopes

- characteristics of telescopes
  - refracting vs. reflecting telescopes
  - collecting area
  - resolution or 'sharpness'
- space-based observatories and other telescopes



Radio Telescope



Optical Telescope

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Two major types of telescopes:

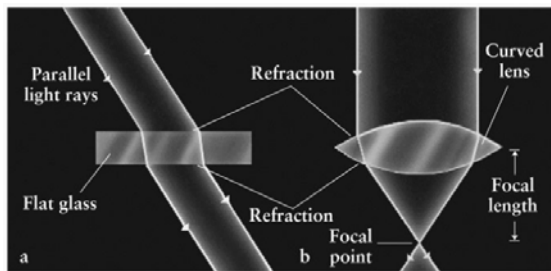
Refracting Telescope



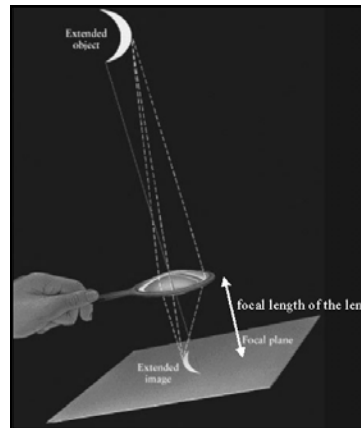
Reflecting Telescope

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A refracting telescope uses a lens to concentrate incoming light at a focus.



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A lens creates an extended image of an extended object.

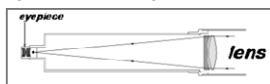
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## Refracting Telescopes

→ use lenses to BEND and FOCUS light



- the first telescopes were refractors
- bigger lenses collect more light, but need a LONGER focal length!
- Lenses eventually became too heavy for their size – began to sag under their weight!



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- famous refracting telescope
- tube is 64 feet long!
- lens diameter = 40 inches or ~ 1 m



Yerkes Observatory near Chicago

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## Refracting Telescopes:

### Advantages:

- nice images after correction for aberration
- lenses do not deteriorate since the tube is sealed
- very rugged, so ideal for amateur work
- inexpensive



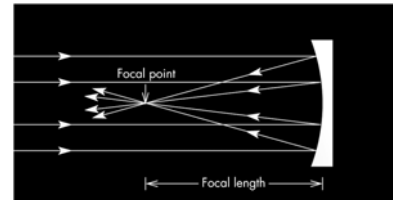
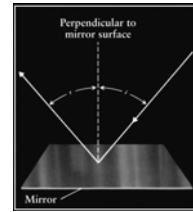
### Disadvantages:

- most glass absorbs UV radiation! only good for redder objects
- suffers from *chromatic aberration* (focus different for all colors)
- difficult to make large, flawless lenses (bubbles in the glass)
- very heavy and must be suspended by its edge (can sag)

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## Reflecting Telescopes:

another way to collect light: use mirrors



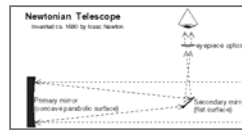
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## Reflecting Telescopes:

All modern research telescopes are reflectors.

### Advantages:

- only the front surface of the mirror must be flawless
- mirrors are lighter and can be supported from behind
- no chromatic aberration
- no UV absorption
- shorter tube



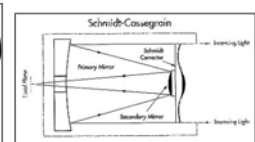
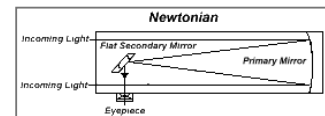
### Disadvantages:

- spherical aberration, if mirror is imperfect
- reflective coating may need to be replaced often
- can be knocked out of alignment easily; not very rugged

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## Different types of Reflecting Telescopes

- Newtonian
- Cassegrain
- others



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## Palomar 200-inch Telescope

largest telescope in the world for almost 50 years

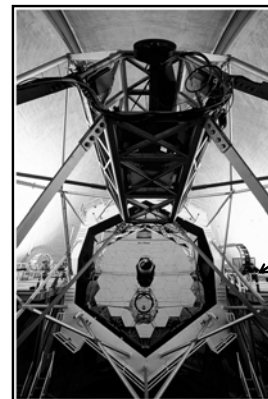


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Keck Telescope

10 m mirror--  
made of  
36 smaller  
parts

Each mirror  
computer  
controlled!



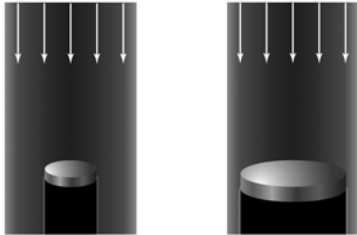
note  
size of  
person!

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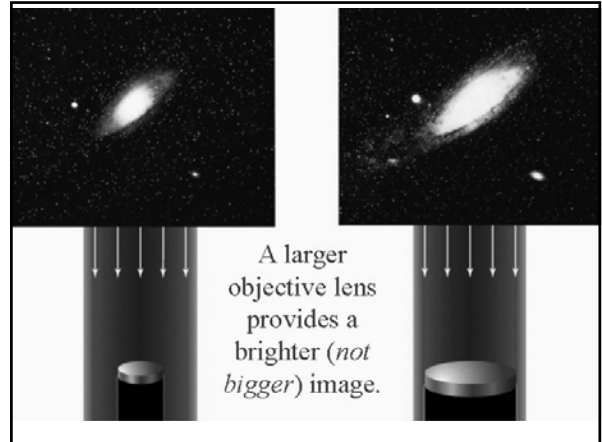
Common features of all telescopes:

1. Size

- *The bigger the telescope mirror or lens, the more light it collects*
- *not to be confused with magnification*

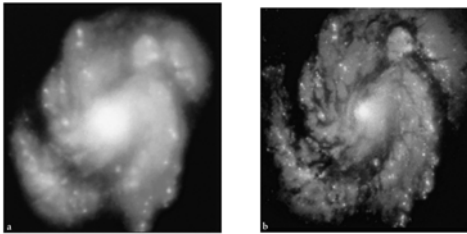


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2. Sharpness/Detail – also known as “resolution”

- *depends on wavelength & diameter of telescope*
- *resolution  $\sim \lambda/D$*

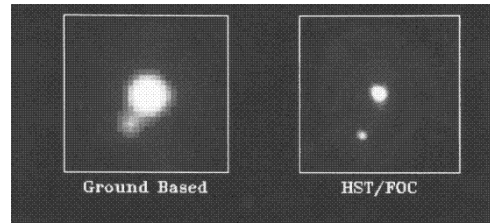


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3. Magnification – not the most important factor for a telescope, but useful

- ***don't buy a telescope based on magnification!!***

$$mag_{telescope} = (focal\ length)_{obj} / (focal\ length)_{eye}$$

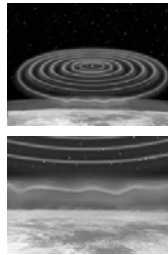


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Earth's atmosphere: observing from the bottom of a swimming pool

Earth's atmosphere distorts images, making them blurry  
And “twinkle”

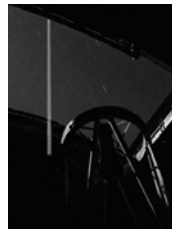
How do astronomers get around this?



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Two modern ways to “beat the atmosphere”

Put your telescope in space

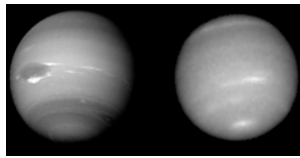


Adaptive optics – bend the mirror to correct for the atmosphere

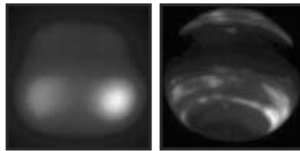
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Neptune with Hubble and Adaptive Optics

Hubble image



Ground-based normal and ground-based with adaptive optics

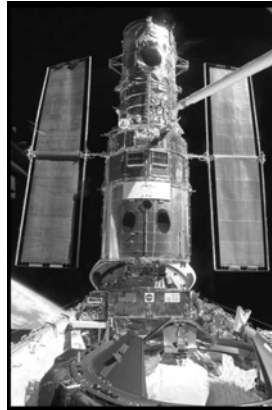


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Space Observatories:

Hubble Space Telescope

- optical telescope
- 2.4-m mirror
- in “low-earth” orbit: 368 miles above Earth
- orbits every 90 minutes

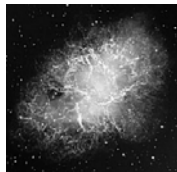


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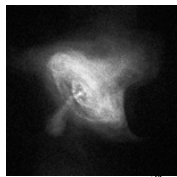
X-rays: Chandra X-ray Observatory



Crab Nebula:



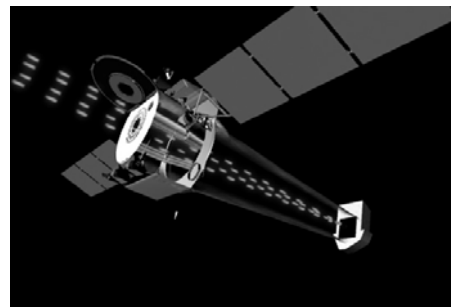
Optical



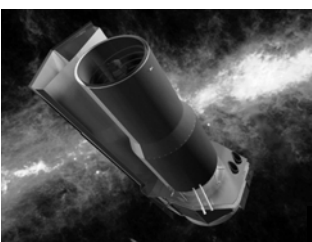
X-Ray

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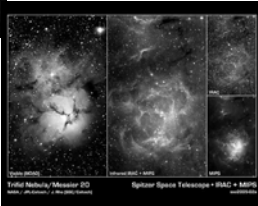
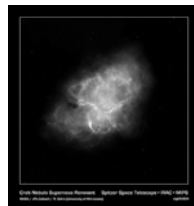
X-rays will pass through normal mirrors to focus X-rays, use “grazing incidence”



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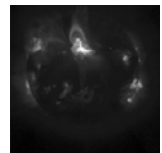
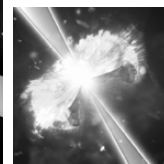
Spitzer Space Telescope  
(in earth orbit: an “infrared Hubble”)



GLAST (Gamma-ray Large Area Space Telescope)

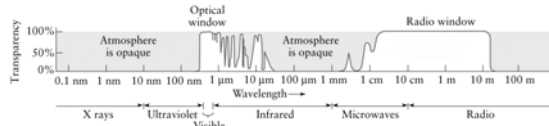


To be launched in 2007



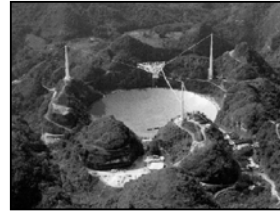
## Other wavelengths from Earth: Radio Telescopes

They are reflecting telescopes that work in the radio region of the electromagnetic spectrum



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## Radio Telescopes

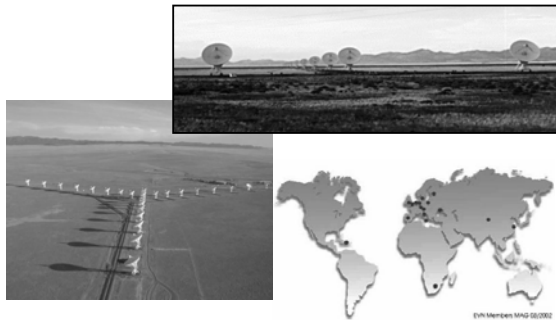


Can have much larger collecting areas than optical telescopes

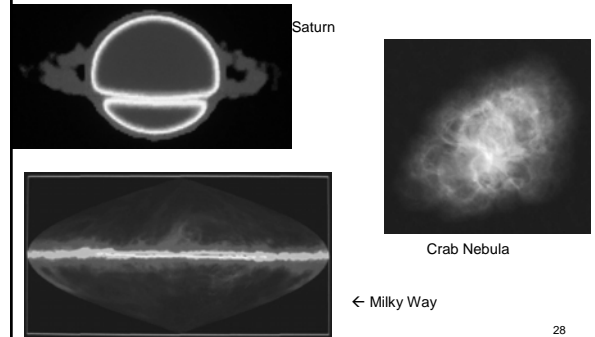


Radio wavelengths are much longer, so you can have "holes" in your telescope →

Radio telescopes can be linked together to form networks that act as a much larger telescope ("interferometry")



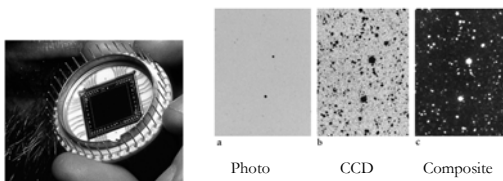
## Radio "images" look different



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## Instruments and Detectors

- The current detector of choice is the CCD
  - CCD = Charge Coupled Device
  - Array of silicon diodes which are light sensitive
  - Same device as used in digital cameras and video cameras.

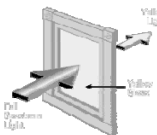


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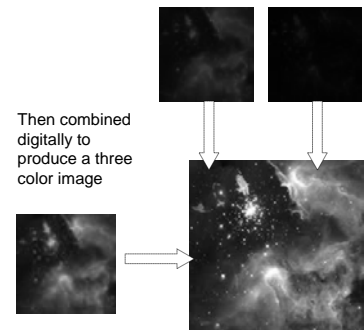
## Three color imaging in Astronomy

CCDs only produce black-and-white images – so how do you get color?

Three or more images of each object are taken, one in each color

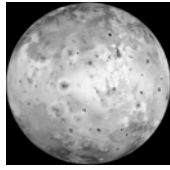


Then combined digitally to produce a three color image



### Examples of color images in astronomy

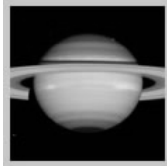
True color



False color

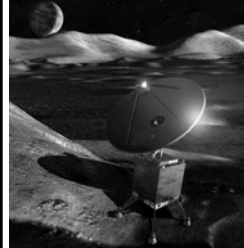
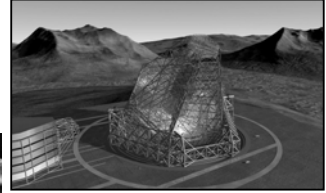


Enhanced color



### The Future of Telescopes

Bigger telescopes:  
OWL – 100 m  
adaptive-optics



More telescopes in space

Maybe even  
on the Moon!

